



α CEN A

α CEN B

Centaur

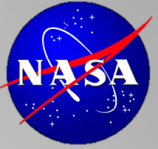
*A scientific and technology pathfinder
for direct imaging exoplanet missions*

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Observatory

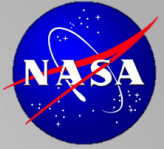


Exoplanet next step: Image an earth-like planet

Centaur enables critical technology for most direct exoplanet imaging

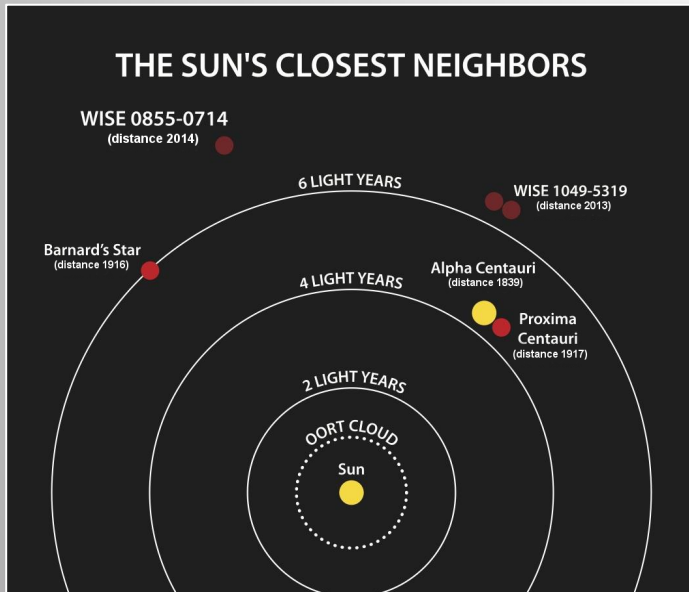
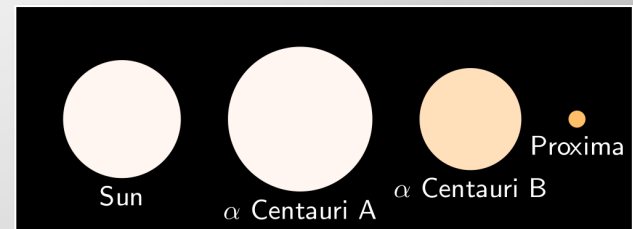
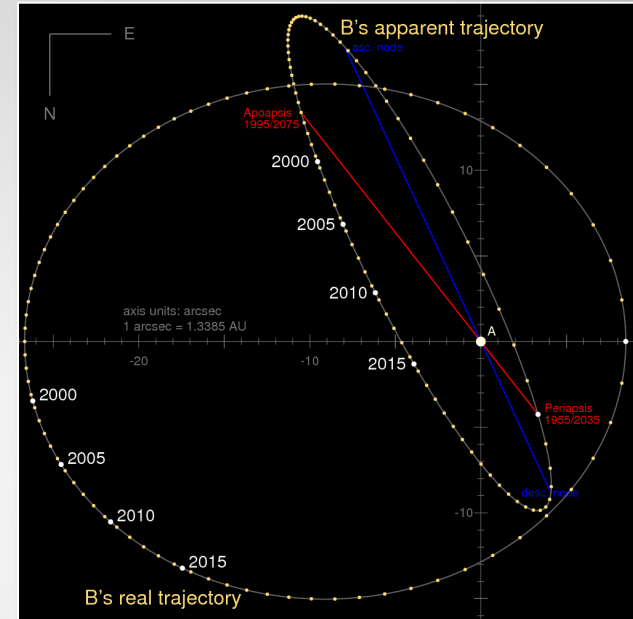
Mission name	JWS I NASA	Centaur Bendek et al.	ACESat Belikov et al.	Ground ELTs TMT/GMT/ESO	WFIRST-AFTA NASA	Exo-C / Exo-S (Stapelfeldt/Seager)
Wavelength/Aperture	IR/6.5m	Vis 0.15m	Vis 0.45m	IR ~ 40m	NIR 2.4m	Vis 1.1m/1.5m
Launch / first light date	2018	2018	2020	2022+	2024	2024
Cost	~\$8.800M	\$10M	\$175M	\$1,000M+	~\$2,000M	~\$1,000
Status	under construction	this proposal	SMEX, submitted	under construction	proposed (Astro 2010 top priority)	study
Detection type	transit and Direct imaging	direct imaging	direct imaging	direct imaging	direct imaging	direct imaging
Planets around Alpha Centauri?	no	Possible	yes	no	possible	possible
Exo-zodi around nearby stars?	Only IR	Possible (>10 Zodi)	Yes	M-dwarf only	Yes	Yes
Require unproven imaging technology?		yes	Yes	Yes	Yes	Yes
Technology demonstration mission?	no	yes	Partially	no	no	No

α Cen AB: a Unique Opportunity for small optical space telescopes



Why Alpha Centauri?

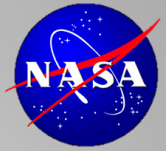
- Alpha Centauri is our closest star and the only one accessible where the Habitable Zone is accessible to a 30cm class telescope
- The system is binary and therefore it double the probability of finding a earth like planet reaching close to 50% chances according to latest Kepler statistics.
- An earth –size planet has been found in 2012, aCen Bb, but is too close to the star. This increases the likelihood of a earth-like planet in the HZ of the star.



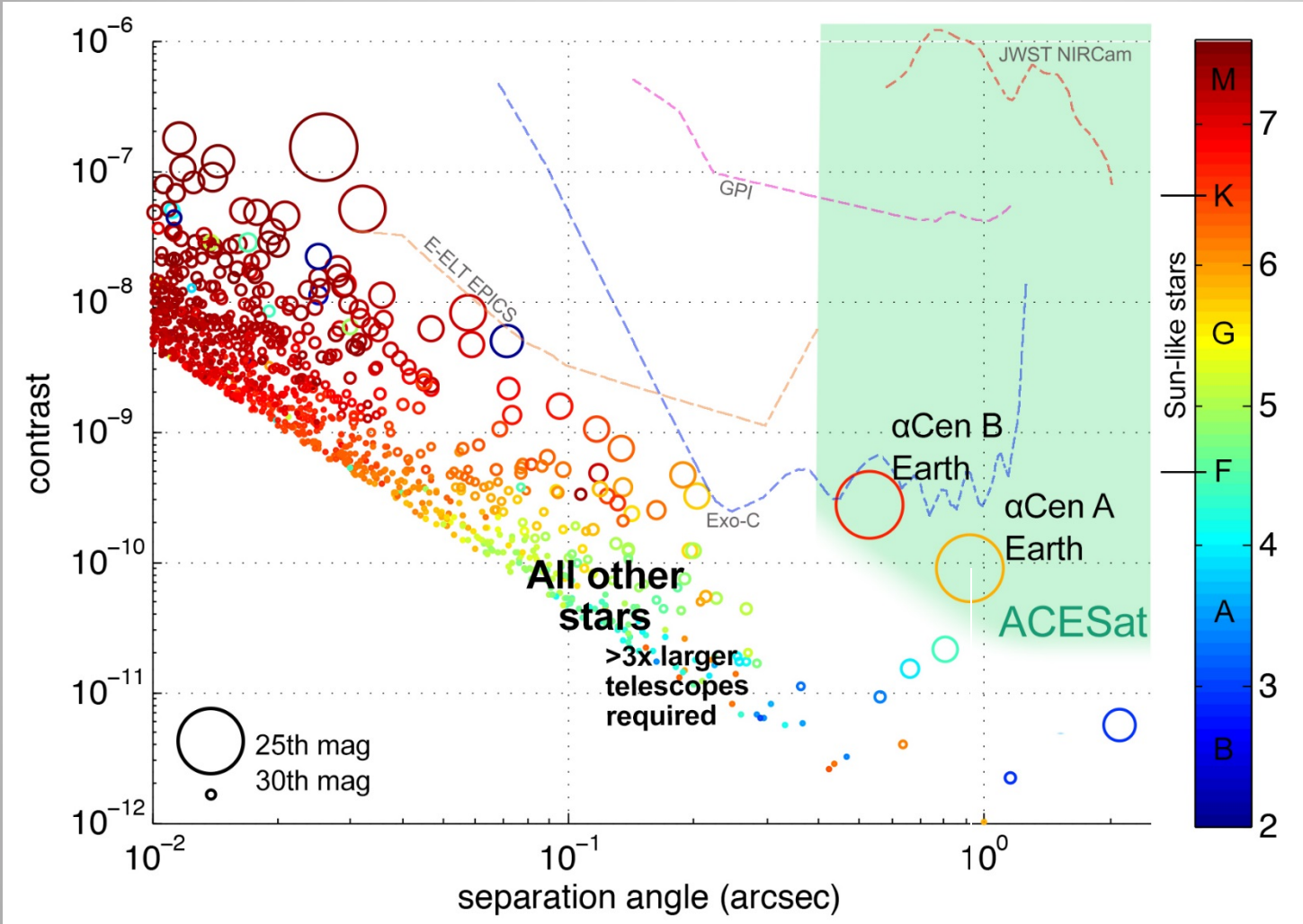
Other science cases

- ACESAT will be also able to measure the exozodiacal light at Alpha Centauri and some other nearby stars. This is critical for other NASA mission design.

α Cen AB: a Unique Opportunity for small optical space telescopes



Simulation of a (hypothetical) Earth twin at quadrature around every nearby star



- Example: α CenA Earth twin with a 30cm telescope at 500nm:
 - separation: $0.92'' = 2.7 \lambda/D$
 - flux: ~ 1 photon per minute for $\sim 10\%$ end-to-end QE (roughly same as for flagship telescope looking at Earths 10pc away)
- α Cen is in a class of its own: any other star requires a $>3x$ larger ($> 10x$ more expensive) telescope

$2.7 \lambda/D$ for 30cm telescope

Value Proposition: Enabling direct imaging of exoplanets around nearby stars leveraging ARC experience and expertise in small sats and exoplanet detection.

- Technology and scientific pathfinder for critical direct imaging technologies.
- Constrain the exozodiacal light on α Cen A & B, ϵ Eri, and τ Ceti. retiring the “dust” risk.
- Detect visible counter part of Radial Velocity planets at ϵ Eri.
- Detection of Neptune like planets around A&B Cen.

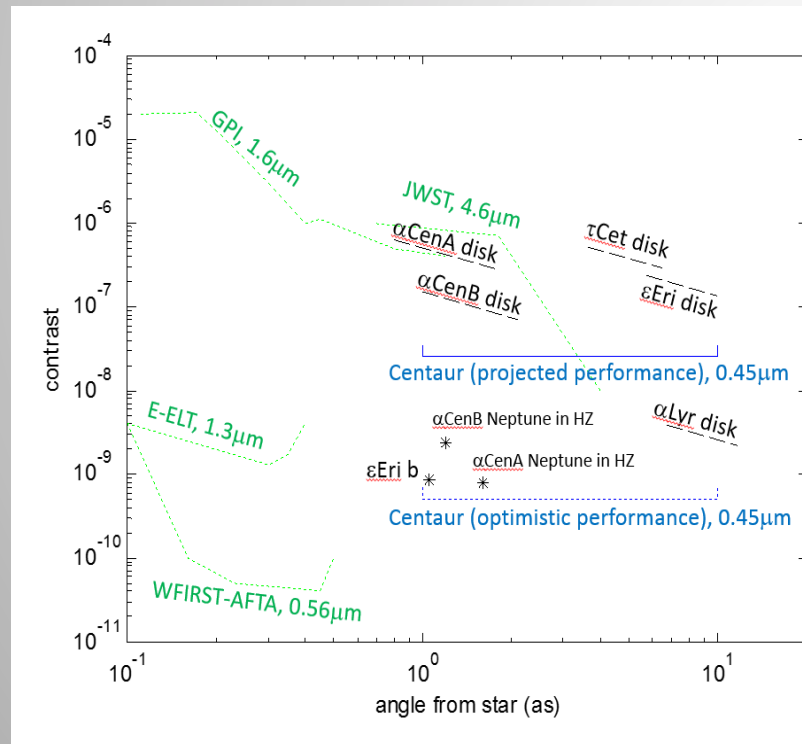


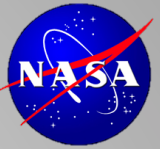
Table 1-2. Disk Science Around Nearby Stars

Object	D [pc]	Disk /Planet Separation		Contrast (based on 100 zodis) ($\times 10^{-8}$)	Integration Times	
		[AU]	[as]		Photon Noise SNR=5 [hrs]	Post-Processing Calibration [days]
α CenA	1.3	0.7-2.8	1-2.1	10	24.5	60
α CenB	1.3	0.6-2.5	0.8-1.8	30	17.2	30
ϵ Eri	3.2	20	6.2	15	2.1	3.53
τ Ceti	3.6	35	9.6	24	11.8	20.6

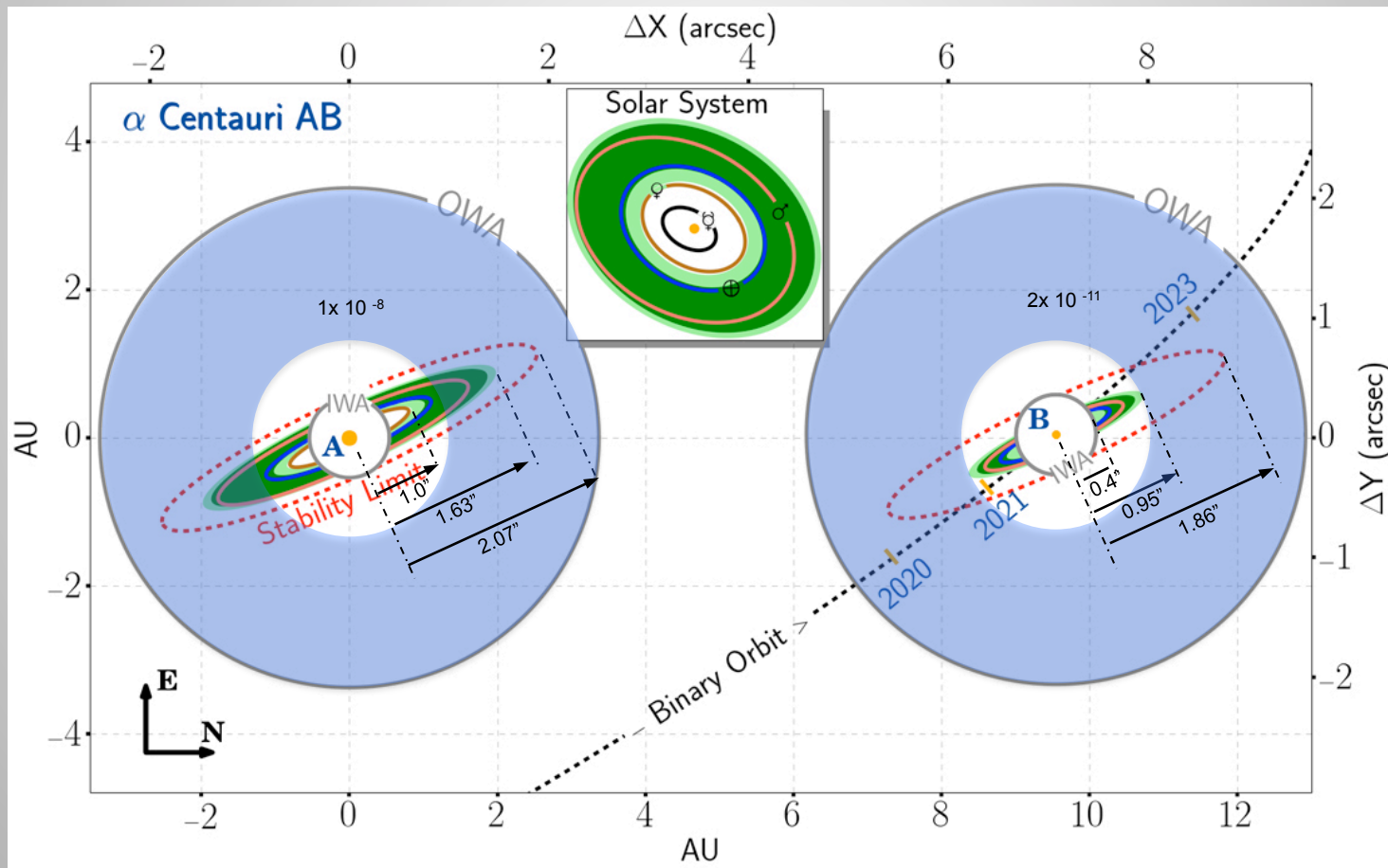
Mission overview

- 1 year, LEO SS, \$10M, Launch Jan 2018
- Aperture 15cm @450nm
- 1 band 15% bandwidth
- IWA @1.6L/D = 1.0"
- OWA @3.5L/D = 2.5"
- 1×10^{-7} raw contrast
- Distance α Cen A and B by 2018 = 4.6" or 7.5L/D

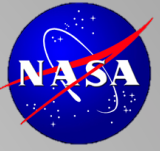
aCen A&B Science



Capable of imaging most aCen A HZ and part of aCen B to constrain down to 30 Exo-Zodi or detect Neptune-like planets.

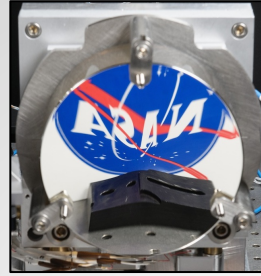
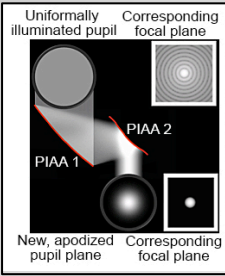
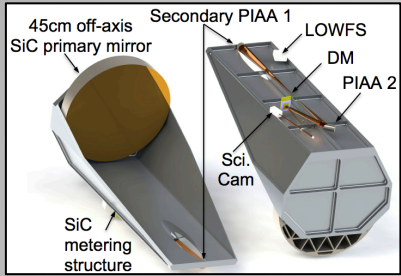


Credit: Billy Quarles, NASA Ames



Instrument Building blocks

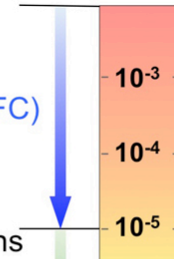
45 cm off-axis telescope with an **embedded PIAA** -> 10^{-5} (1.6 – 10M/D)

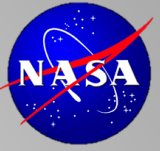


Diffraction from on-axis star (no coronagraph)

Coronagraph (w/o WFC)

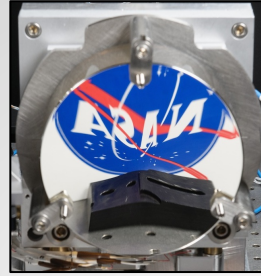
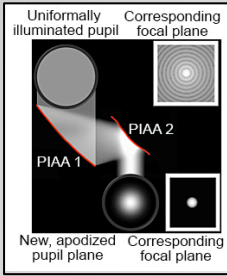
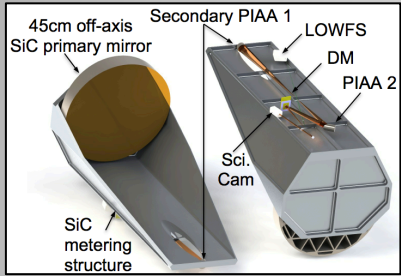
Aberrations from on-axis star; diffraction + aberrations



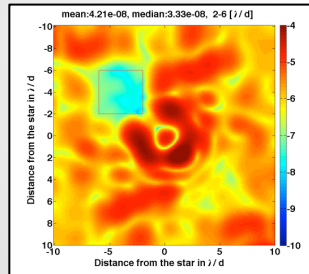
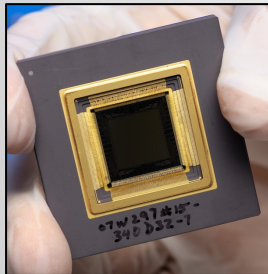
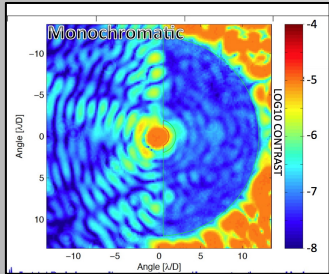


Instrument Building blocks

45 cm off-axis telescope with an **embedded PIAA** -> 10^{-5} (1.6 – 10M/D)



WFC (Multi-Star Wave Front Control) -> 10^{-8}



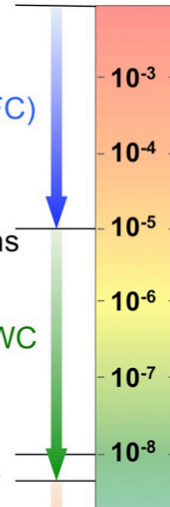
Diffraction from on-axis star (no coronagraph)

Coronagraph (w/o WFC)

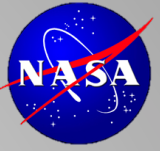
Aberrations from on-axis star; diffraction + aberrations from off-axis star

MSWC

Raw noise floor after MSWC *required* *expected*

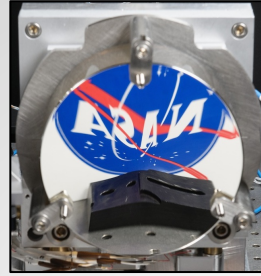
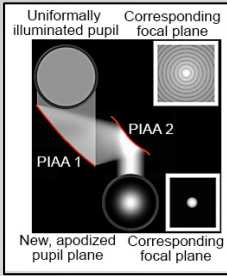
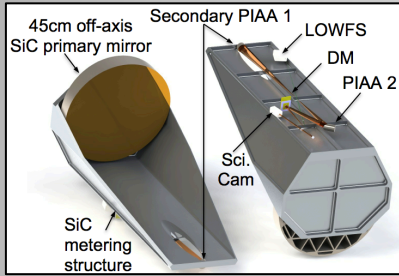


100 zodis (60° inc)

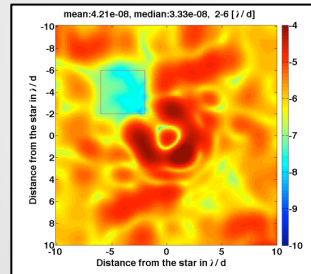
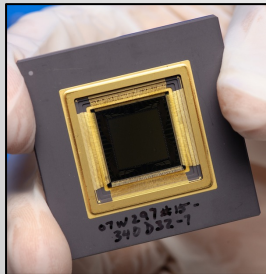
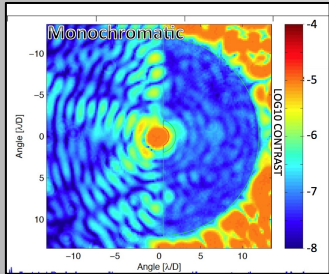


Instrument Building blocks

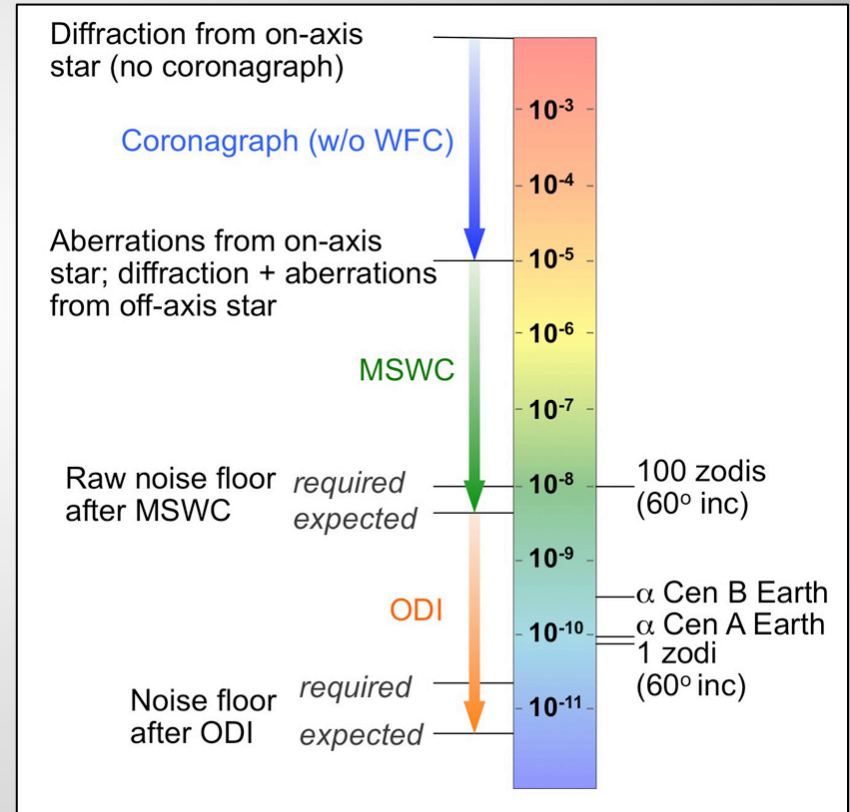
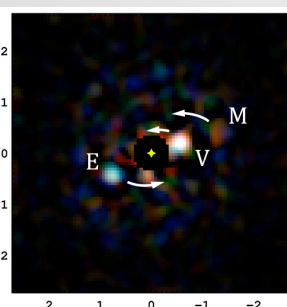
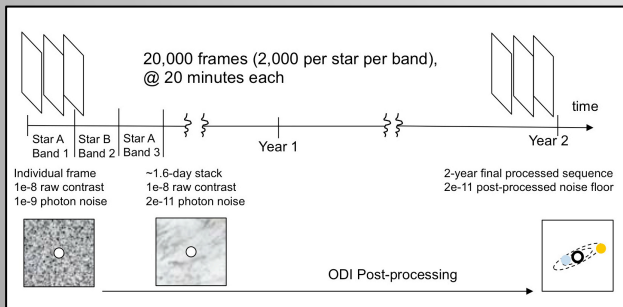
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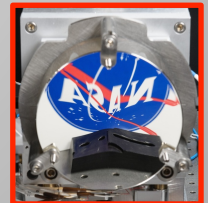
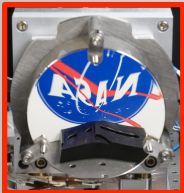
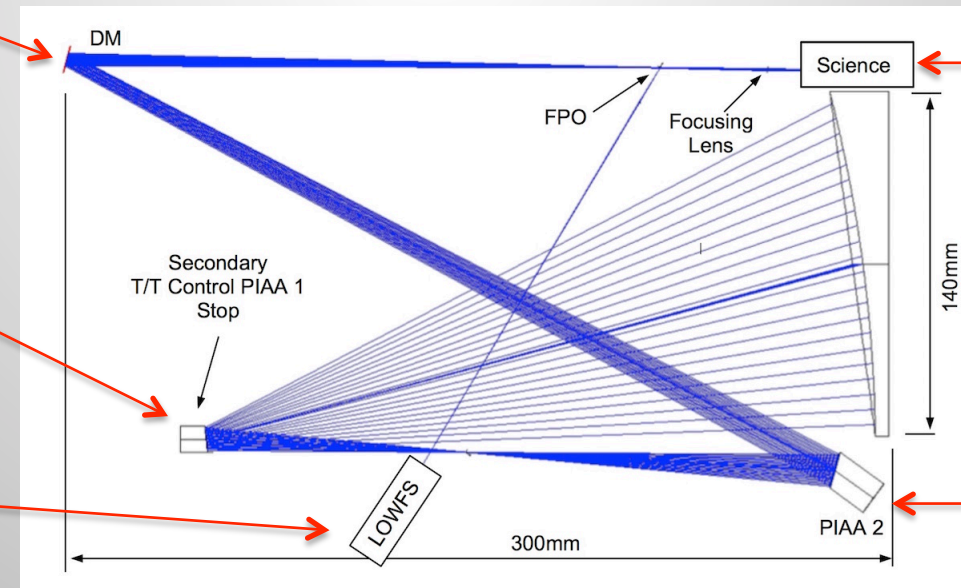
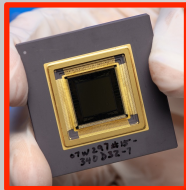
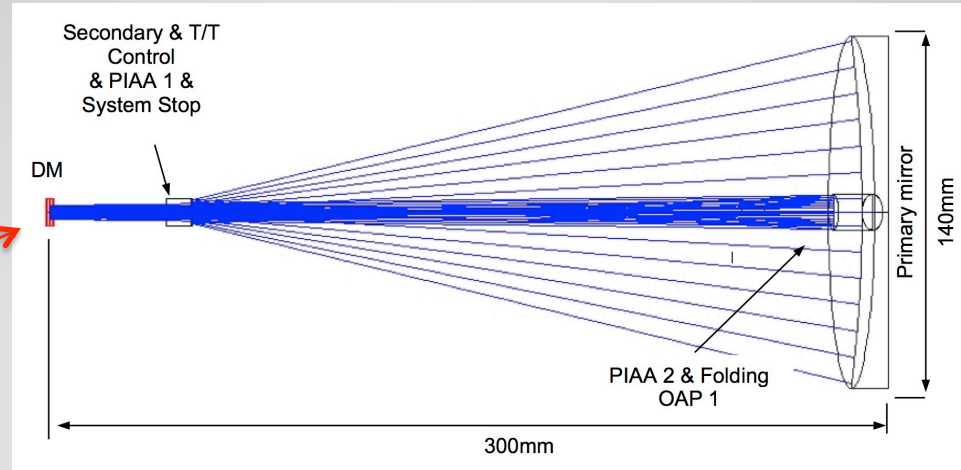
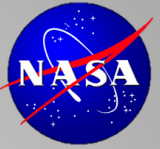
WFC (Multi-Star Wave Front Control) -> 10^{-8}



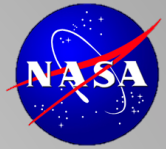
Continuous observation ODI -> 10^{-11}



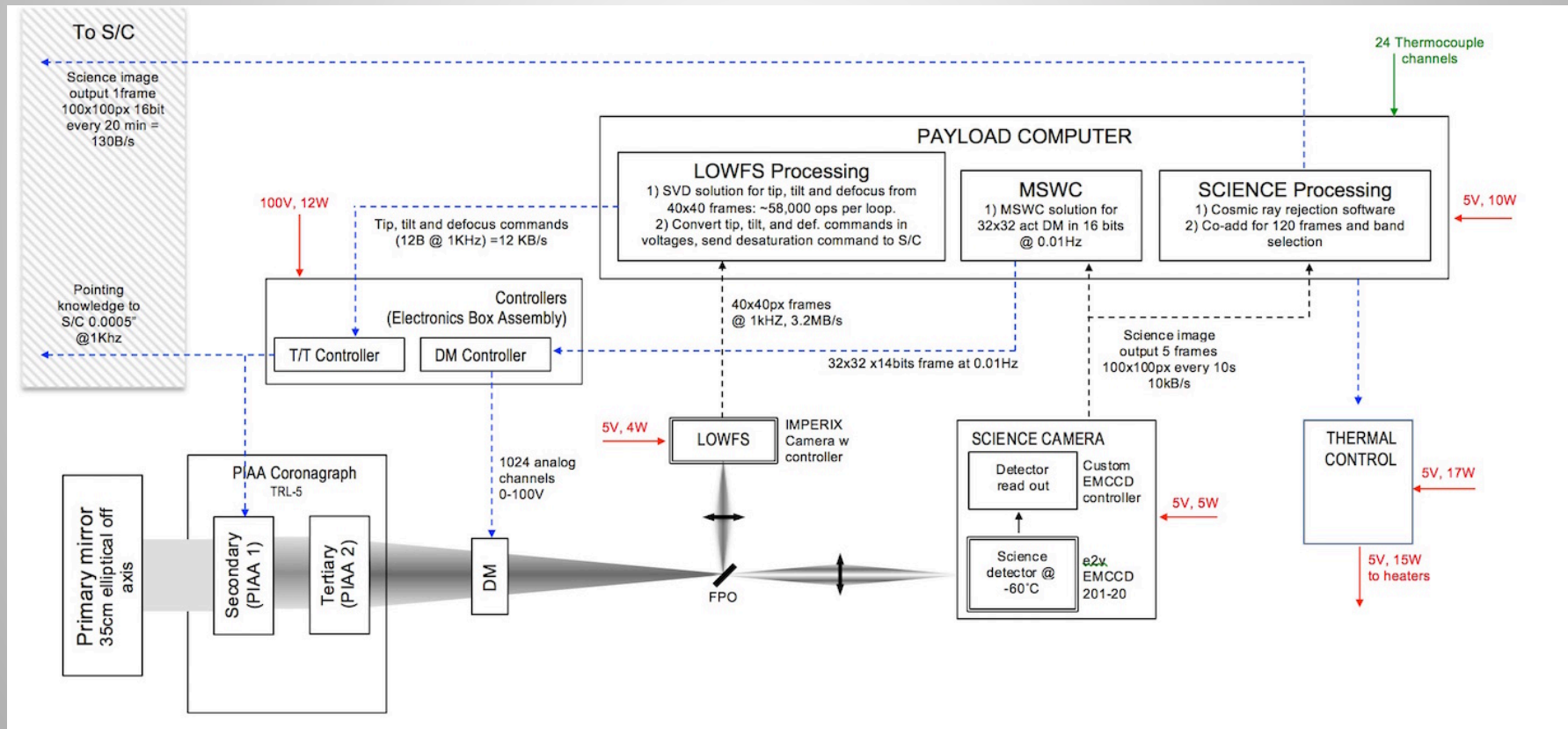
Centaur Optical Design

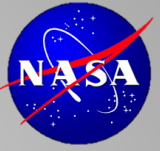


Centaur Functional diagram



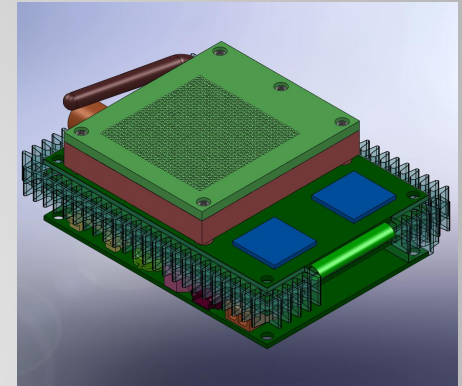
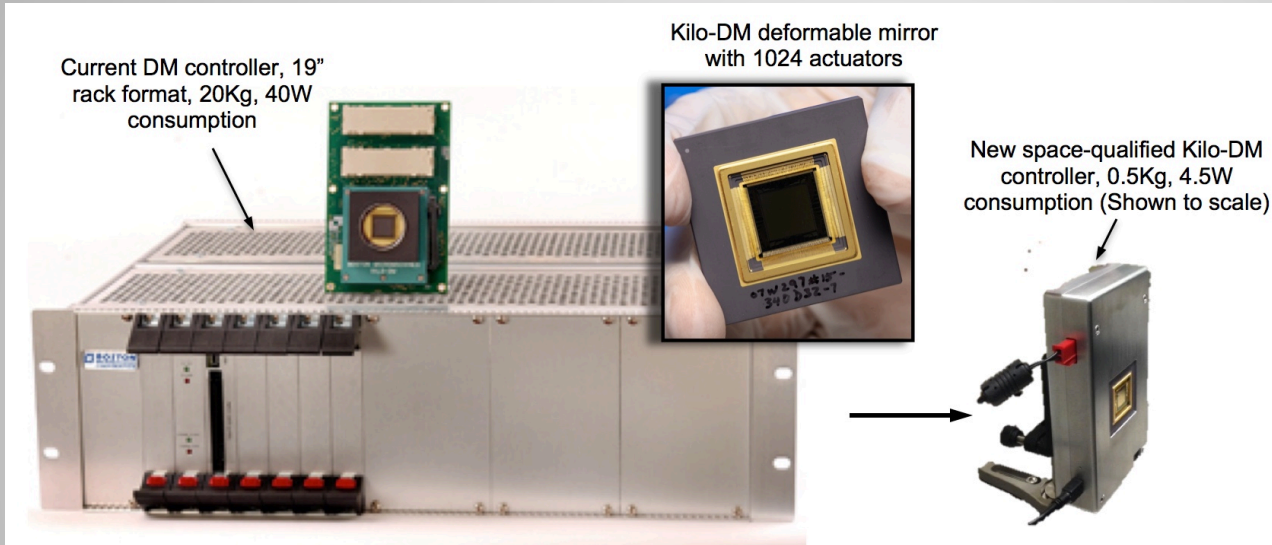
- Centaur has the stop located at the secondary mirror which is also PIAA 1 and tip-tilt mirror
- The tip-tilt mirror is controller by the knowledge provided by the LOWFS
- For fast Jitter control and small amplitude we use the DM to control



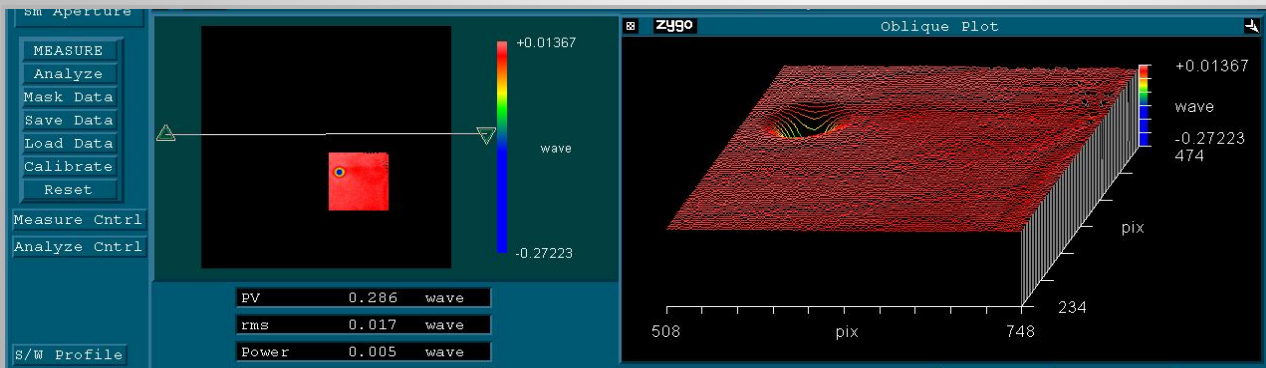


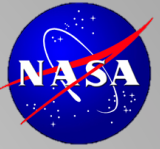
DM Controller

- New compact Kilo-DM (1024 actuators) controller being developed for this mission



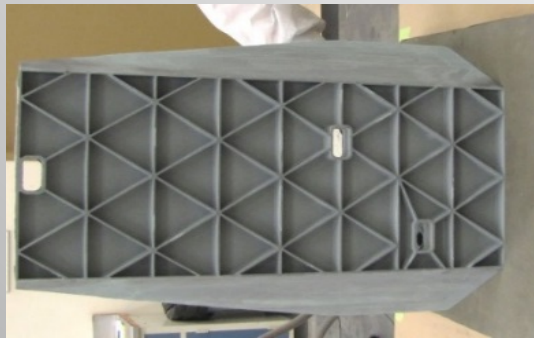
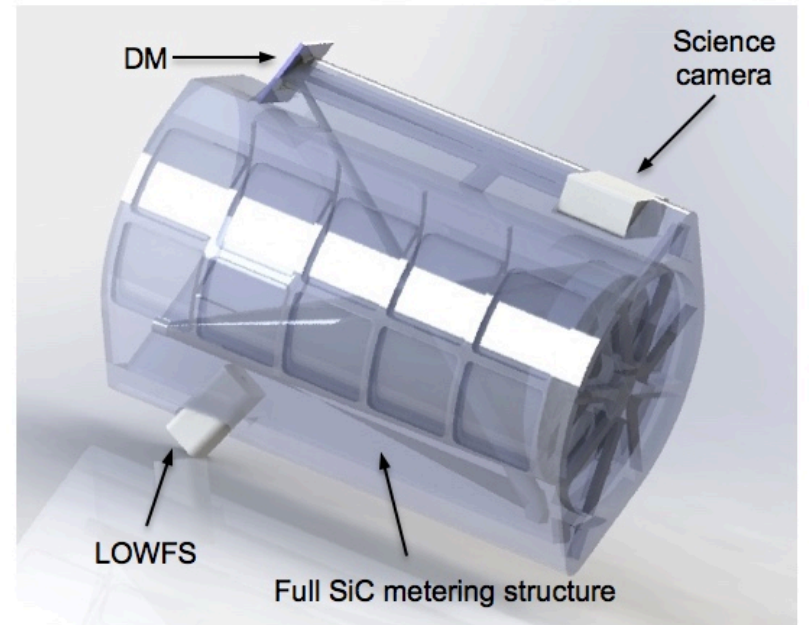
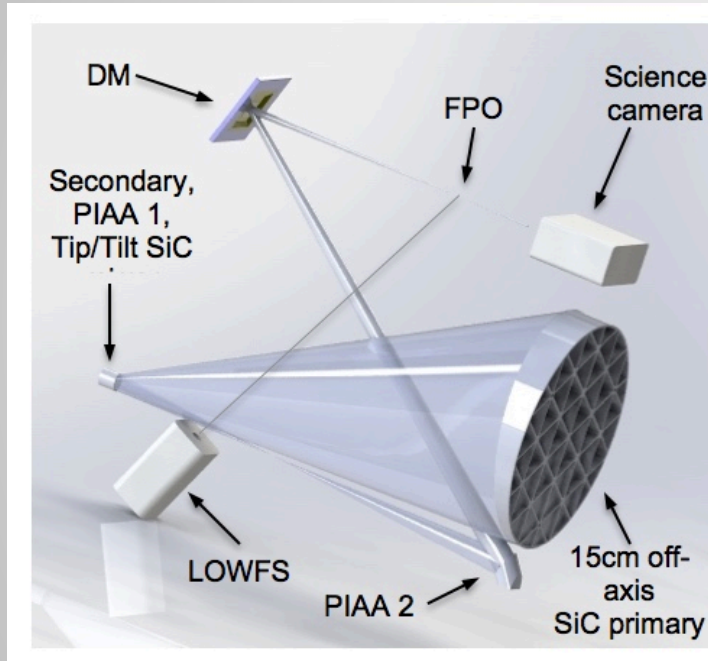
- Model of Kilo-DM space qualified controller currently under design



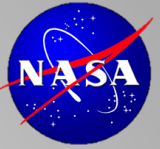


Telescope Hardware

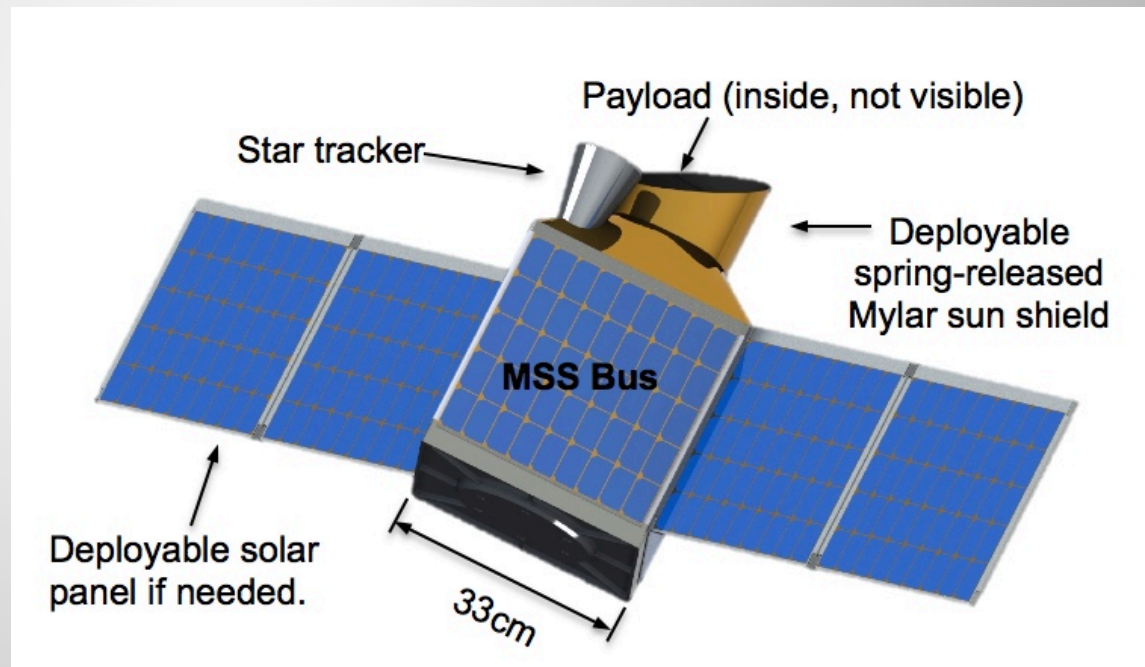
- Full SiC 15cm, Off-axis telescope, L/25 max end-to-end WFE (Total 15Kg mass)
- Active thermal control to maintain 10°C operation with 0.3°C PV stability
- 1.5mas RMS stability LOWFS (Demonstrated for CAT III EXCEDE Lockheed Martin)



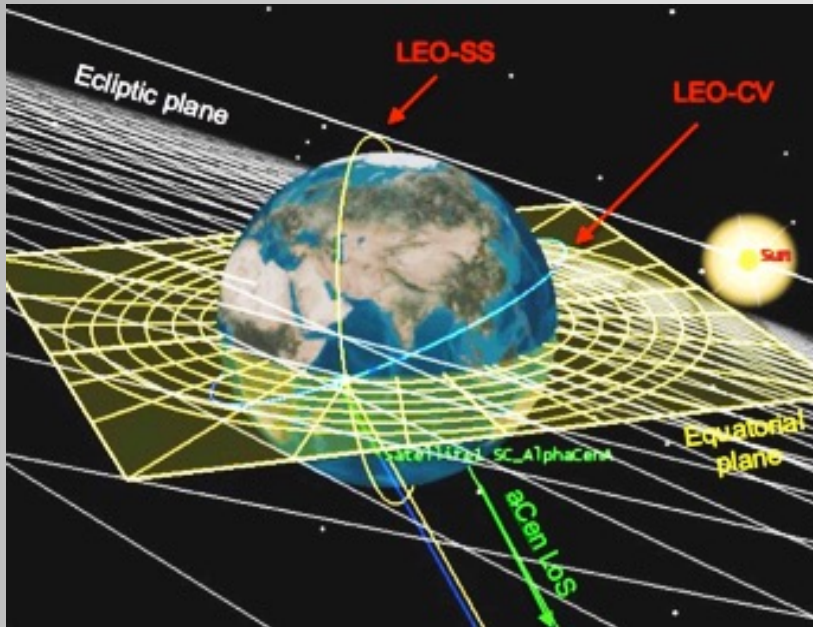
Spacecraft



- Low Cost off-the-shelf Millennium Space System 27U bus (30x30x30cm)
- We will buy one unit “as is” and will improve stability.
- Low data rates requirements allow to use normal ground stations
- 800km orbit avoid decay problems and increases target access.



Mission operations



Low Earth Orbit concept

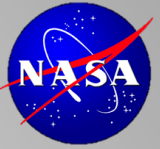
Resilient to orbit inclination

Baseline of 800km Sun Synch.

- 23° inclination would allow ~25day of uninterrupted view of aCen every 3 months.
- Sun Synch provides thermal stability.
- Launch on Falcon 9, utilizing TriSept utilizing FANTM-RiDE

Table 2.2-1. Alpha Centauri (left) and Epsilon Eridani (right) Trade Inclination Vs. RAAN (Alt 800 km)

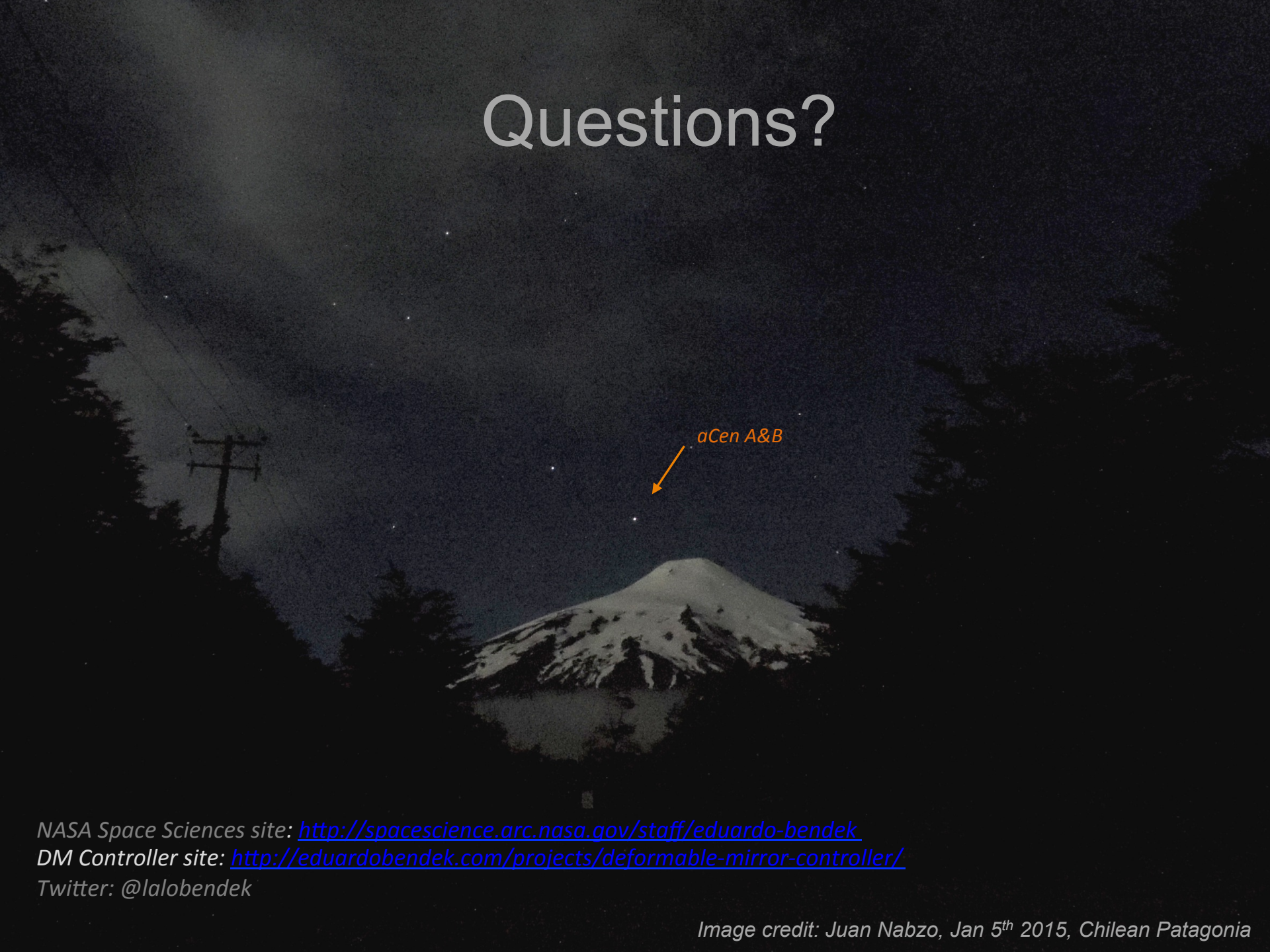
Alpha Centauri	Best Case	Worse Case		Epsilon Eridani	Best Case	Worse Case
Inclination (deg)	90	0		Inclination (deg)	10	90
RAAN (deg)	120-160, 300-340	0-360		RAAN (deg)	0-360	0-360
Total Access (days)	365	240		Total Access (days)	329	240
Max Cont. Access (days)	365	0.05		Max Continuous Access (days)	22	0.05
Min Cont. Access (days)	8760	1.1		Min Continuous Access (days)	1.3	1.1
Mean Access Per Orbit (%)	100	65		Mean Access Per Orbit (%)	91	65



Conclusion

- 1) We developed an instrument design that can measure ExoZodi around nearby stars.
- 2) We developed a mission concept in low earth orbit utilizing a 27U off-the-shelf microsat
- 3) We are advancing key technologies (PIAA, DM, WFC, Post-processing) for ACESat and other direct imaging missions (AFTA-C, EXO-C, EXCEDE)
- 4) Mission is expandable to *Centaur "plus"*, **Capable of imaging earth like planets** with a 25 cm telescope that **will fit on the same bus**

Questions?



NASA Space Sciences site: <http://spacescience.arc.nasa.gov/staff/eduardo-bendek>

DM Controller site: <http://eduardobendek.com/projects/deformable-mirror-controller/>

Twitter: @lalobendek

Image credit: Juan Nabzo, Jan 5th 2015, Chilean Patagonia